LAD 14 DIAMOND LIGHT SOURCE

Dave Bond





So, what do we actually do?

- The Diamond machine is a type of particle accelerator
- CERN = high energy particles smashed together and analyse the "crash"!
- Diamond = accelerate electrons to produce synchrotron light
- Use this light to study matter like a "super microscope"



A Brief History of Synchrotrons

- Early particle accelerators produced "waste" in the form of light
- However X-rays were increasingly used as diagnostic tools, such as crystallography.
- 1953 Structure of DNA realised using X-ray crystallography
- 1956 first "parasitic" experiments carried out at a synchrotron









A Brief History of Synchrotrons

- 1960s Synchrotron light used for atomic and molecular spectroscopy
- 1970s increasing number of SR experiments, including crystallography
- 1980 first dedicated synchrotron light source was built in 1980 in the UK - SRS
- Diamond has replaced the SRS which closed in August 2008



The Diamond machine

- Three particle accelerators:
- Linear accelerator
- Booster Synchrotron
- Storage ring
- (48 straight sections angled together, 562m long)









It's all done with magnets

We use magnets to focus and direct the beam, and to make the light we use in experiments



B = 1.4 T (20,000 x Earth's magnetic field, or 100 x typical bar magnet).

24 straight sections in the storage ring. 22 straights available for insertion devices (IDs)

These enable us to produce X-rays that are high energy, more tightly focussed and tuneable.



What is synchrotron light?

The many **COLOURS** of light



Types of experiment



Research into and much more











Casting aluminium

Understanding the corrosion process

Pharmaceutical manufacture and processing Organic photovoltaics Tunable polymers



MOFs for hydrogen storage



Bio-mimetics



Multiferroics – electronic storage and memory



Harry's wheel -

complex templates



Earth science



Environmental science



Cultural

conservation

aritage and



research



Understand rejection in hip implants





What lustre setup do we have?







10G Ethernet

•4 OSS servers

•1.2TB MDS storage connected via FC. DDN supplied EF3015

• 470 TB OST storage in 30 volumes spread over 4 OSS nodes

•OSS storage IB connected DDN SFA10K 8+2 RAID

0) -			-			-					0
			Þ	5		Þ	-	\mathbf{s}		-		
				フ		F	·3	01	$\mathbf{\overline{b}}$	-		:
				-			-			-		
			\triangleright	-			-			-		
0							-			-		0
0				-		\triangleright	-		\triangleright			0
							-		\triangleright	-		
			\triangleright			\triangleright	-			-		•
				-			-		\triangleright	-		
			\triangleright	-			-		\triangleright	-		
0				-			-			-		0
0							ΞΔ	10	X	1000		0
			F	2	L		4	स्य	Ð	-		
				-			-			-		:
				-		\triangleright	-		\triangleright			
			\triangleright	-			-			-		
P												0
0				-			-			-		0
				-			-			-		
										-		•
							-					
							-			-		
0		(11)			(11)			653			(11)	0



How Lustre at Diamond has evolved....

- The first production systems were Lustre 1.6
 - 10G Ethernet connected
 - DDN 9900
- The second production file system was Lustre 1.8
 - 10G Ethernet connected
 - DDN SFA10K
- Both file systems have been upgraded to at least a couple of versions of Lustre 1.8
- Most recently upgraded to Lustre 2.5 with the second of the two production file systems now is IB capable



What other filesystems do we use?



1.1PB of available storage used for low data rate beam lines and on-line archiving



150TB of available storage used for data not collected at Diamond, home areas, databases, virtual machines



877TB of available storage used for high data rate beam lines



814TB of available storage used for high data rate beam lines



Why do we have GPFS as well as Lustre?

- Single stream performance is better for GPFS, we were aiming for 900Mbs sequential read and write. This was a hard requirement for two beam lines.
- Though we achieved this we have had issues since.
 - Mixing 1G and 10G crippled performance. When writing over 10G and reading over 1G. There was large performance penalties.
 - Upgrading to IB on the NSD servers and communicating to the cluster over IB resolved this
 - Native CIFS access is difficult to manage and deploy, we had regular issues where it would drop out of the GPFS cluster.
 - We are currently using SAMBA for CIFS access and RSYNC as a data mover.
 - In GPFS 3.5 the CIFS issues are believed to have been improved. We are awaiting the DDN release.





Upgrade to 2.5 and what we hope to gain?

We currently have strict archiving rules as we aim to only store 6 months of data on disk, before access is only from tape or for some data archive disk.

HSM -> Change Logs

This will aid us with robin hood and the ability to gather archiving lists easier. HSM may in the future automate moving files to low data rate storage.





Upgrade to IB and what we hope to gain?

IB

After a recent upgrade to IB on our GPFS system this was the logical next step to upgrade Lustre to provide high bandwidth low latency connectivity between the storage and the cluster

6036 Spine Switch 6025 Leaf Switch





Current Diamond Infiniband layout - 09/09/2014









How we went about testing...

- MDTEST
- \${MPI}/bin/mpirun -mca btl self,tcp,sm --hostfile \${UNIQHOSTS} -np \${jobs} ~bnh65367/code/mdtest/mdtest -I 10 -z 5 -b 5 -i 5 -u -d \${TESTDIR}
 - Run with our cluster one process per host
 - -I 10 -> 10 items per directory
 - -z 5 -> Tree depth of 5
 - -b 5 -> branching factor of 5
 - -u -> create unique working directory for each task
 - -d -> working directory
- IOR

\$MPI/bin/mpirun -mca btl self,tcp,sm --hostfile \${HOSTFILE} -np \${jobs}
/home/bnh65367/code/ior/src/ior -o \${TESTDIR}/ior_dat -w -r -k -t1m -b \$
{BLOCKSIZE} -i 5 -e -a POSIX

- -w -> write file
- -r -> read existing file
- -k -> keep the file do not remove it
- -t -> maximum transfer size 1m
- -b -> block size
- -i -> repititions

IOZONE

- -e preform fssync
- -a -> use posix



iozone.x86_64 -i 0 -i 1 -s \$2 -r 4k -t \$3 -+m ~/dls-science-user-area/iozone-hos

MD Test





IOR After code upgrade





MDTEST

Lustre03



This work only has been completed in the last few weeks, testing is still ongoing.

The first users of the upgraded Lustre system will be using it today.



Future HPC filesystems at Diamond?

With new high data rate detectors already being tested and the new national electron microscopy facility being built to hold 4 electron microscopes

The planning for the next HPC file system has begun!



Examples of detectors at DLS

Specification	PCO 4000 PCO 4000 and PCO Edge are high speed cameras developed and supplied by PCO AG,Donaupark 11, 93309 Kelheim, Germany	PCO-Edge	Pilatus 6M Pilatus 6M is a very high capability detector developed and supplied by DECTRIS Ltd. Neuenhoferstrasse 107, 5400 Baden, Switzerland	Excalibur Excalibur is a detector development collaboration between the Science and Technology Facilities Council and Diamond Light Source - Journal of Physics: Conference Series Volume 425 Part 6 J Marchal <i>et al</i> 2013 <i>J. Phys.: Conf. Ser.</i> 425 062003 doi:10.1088/1742-6596/425/6 /062003	Percival PERCIVAL (Pixelated Energy Resolving CMOS Imager, Versatile and Large) is an Ongoing development project between DESY and RAL / STFC
Frame	2D	2D	2D	2 - 3D	2D
Scan Size	1D	1D	1 - 3D	1 - 2D	1D
Frame rate	5Hz	100Hz	100Hz	100Hz	120Hz
Data Rate	100MB/s	700MB/s	~640MB/s	~600MB/s	~5-6 GB/sec
Status	Complete	In development	Complete	Commissioning	In development





At the moment the contenders are GPFS and Lustre Lustre is preferred because of the seemingly high sensitivity of GPFS for issues in a distributed environment

Requirements

- •6 machines writing into the same file at 10GB/s, just for one beam line
- •Saturating 10GB/s with a single stream or an IB connection even better
- •A nice to have would be a native lustre client for Windows







